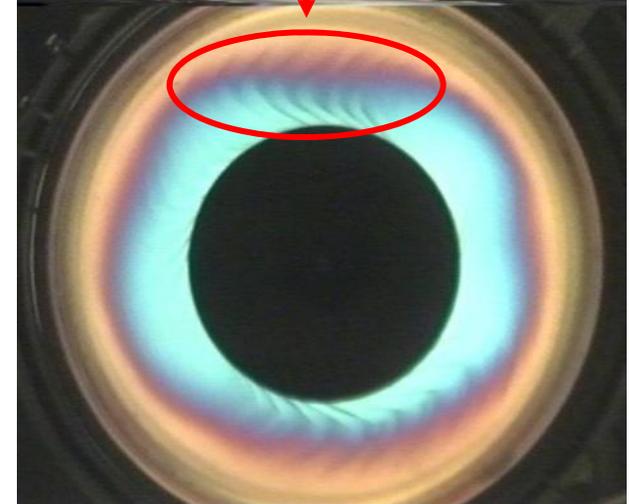
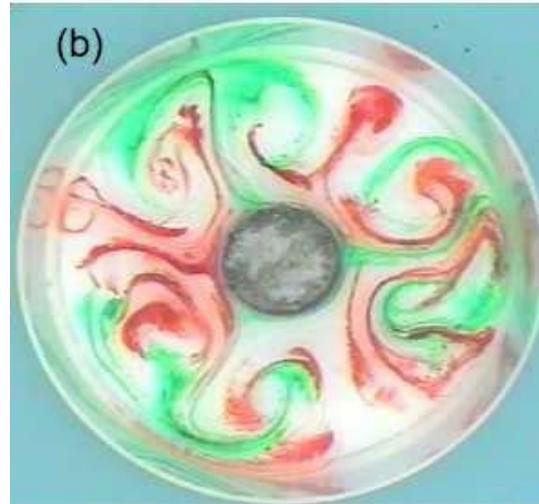
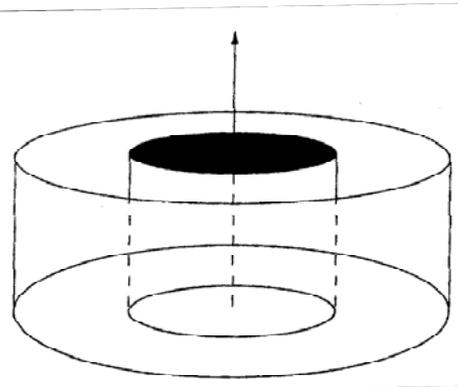
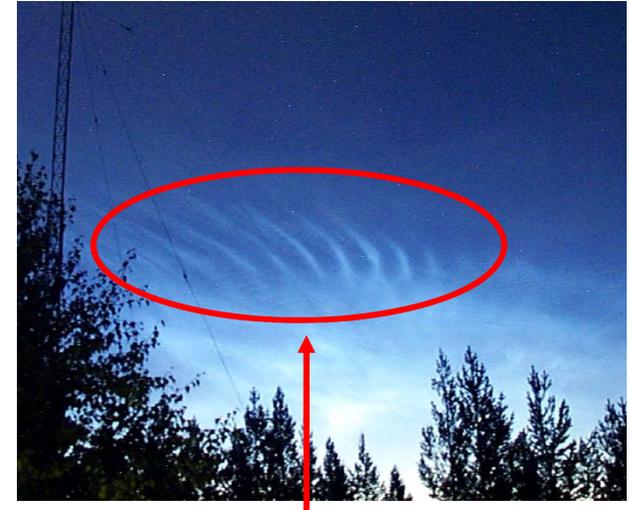
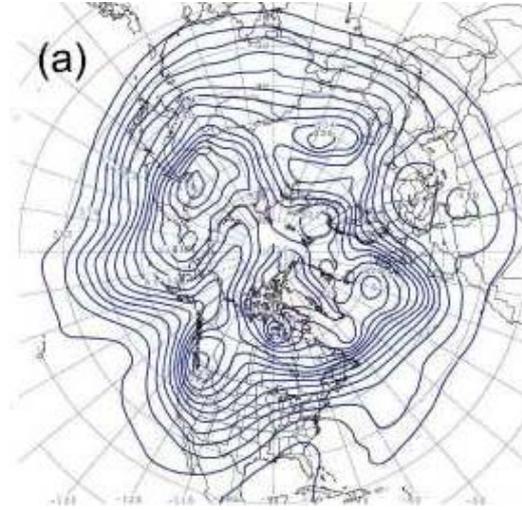
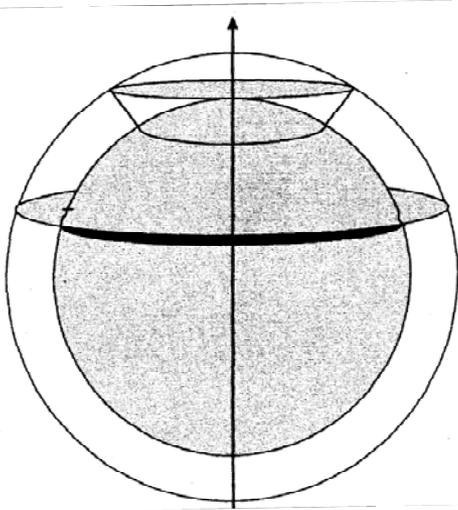


# Gravity Waves in the Laboratory: Mechanisms, Properties, and Impacts

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# Rotating planet –v– rotating annulus



(Read et al. 1998)

(Ravela et al. 2009)

(Williams et al. 2008)

See movie at [www.youtube.com/RotatingAnnulus](http://www.youtube.com/RotatingAnnulus)

Gravity waves  
at the interface  
in a two-layer  
laboratory  
experiment



blue = high

yellow = low



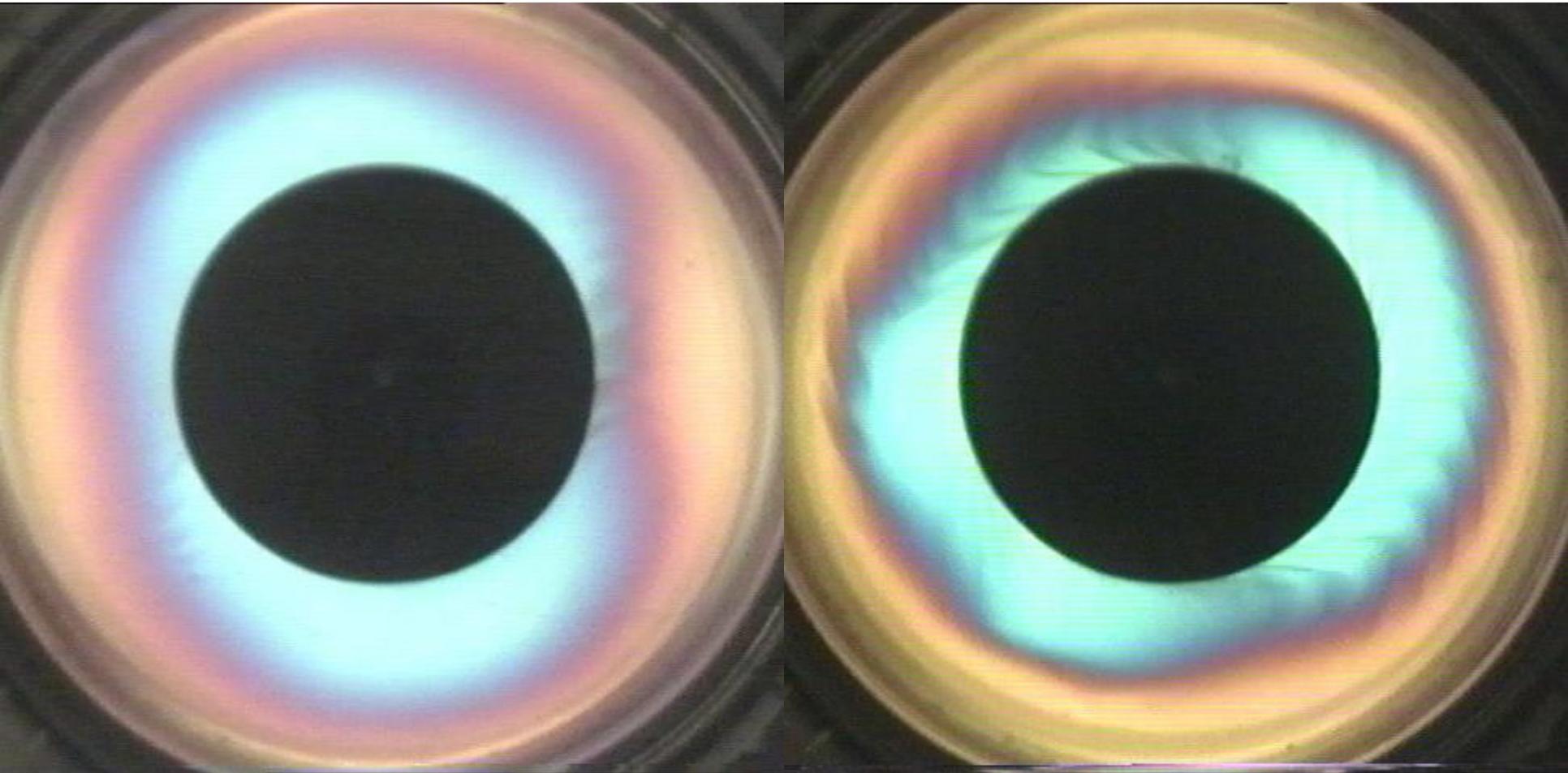
# Generation mechanism

# Possible generation mechanisms for the gravity waves

- Flow over topography
- Convection
- Kelvin-Helmholtz shear instability
- Surface forcing of the ocean by the atmosphere
- Lateral boundary instability
- Geostrophic-adjustment emission from an unbalanced initial state
- Spontaneous-adjustment emission

# Properties

# Variation of gravity wave amplitude with $Ro$

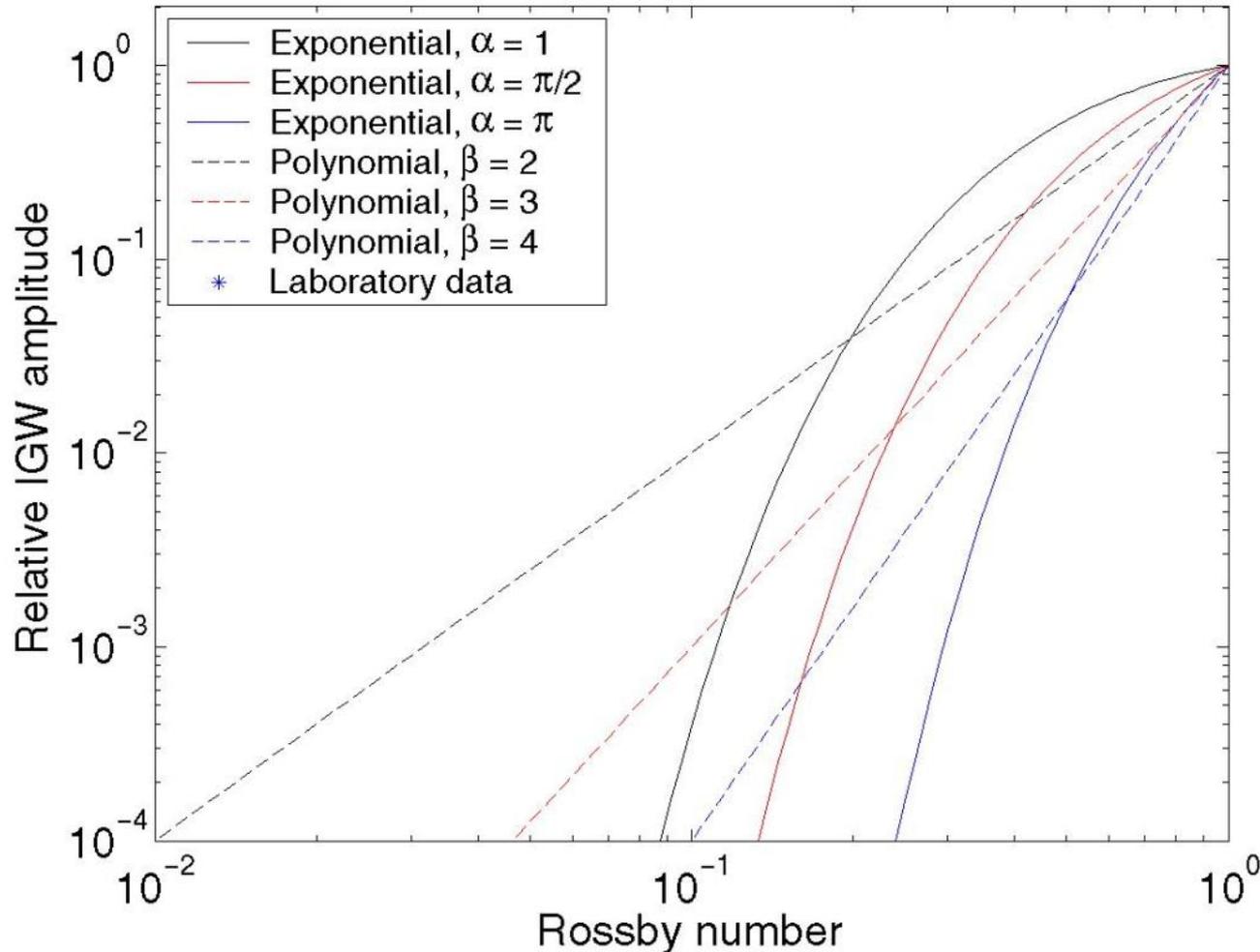


$Ro = 0.05$

$Ro = 0.14$

(Williams et al. 2008)

# Variation of gravity wave amplitude with Ro



- non-asymptotic theories suggest  $\sim Ro^{-1/2} \exp(-\alpha/Ro)$  (e.g. Vanneste & Yavneh 2004; Plougonven et al. 2005 find that  $\alpha \geq \pi/2$ )
- standard asymptotic analysis suggests  $\sim Ro^\beta$ , where  $\beta \geq 2$
- but the laboratory data suggest  $\sim Ro^1$ , **apparently in contradiction to theory**

(Williams et al. 2008)

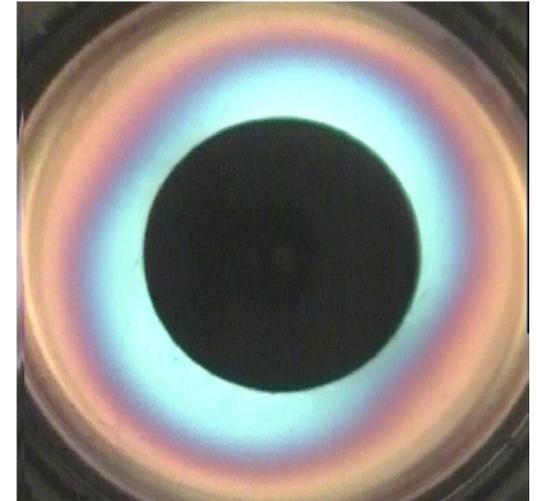
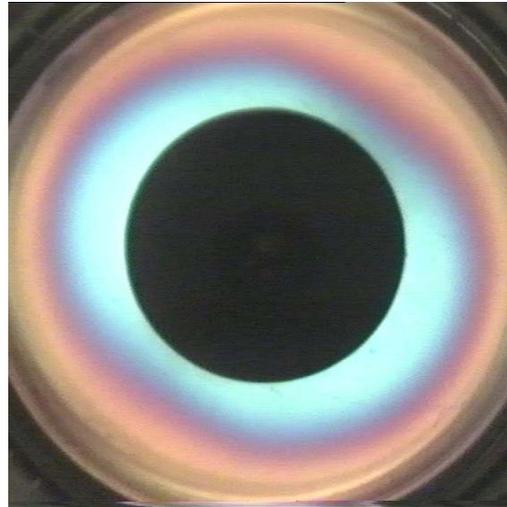
# Wave energetics

- The total energy (i.e. kinetic plus potential) of the fluid has contributions from the waves and the zonal-mean zonal flow.
- For Rossby numbers  $\sim 0.1$ , we find that **91%** of the wave energy is in the large-scale rotational wave, and the remaining **9%** is in the short-scale inertia-gravity waves (Williams et al. 2008).
- These numbers compare well with those from atmospheric reanalysis data (**90%** and **10%**, respectively; Žagar et al. 2009).
- Of the energy in the large-scale wave, we find that  **$\sim 1%$  per 'day'** is fed into the inertia-gravity waves.

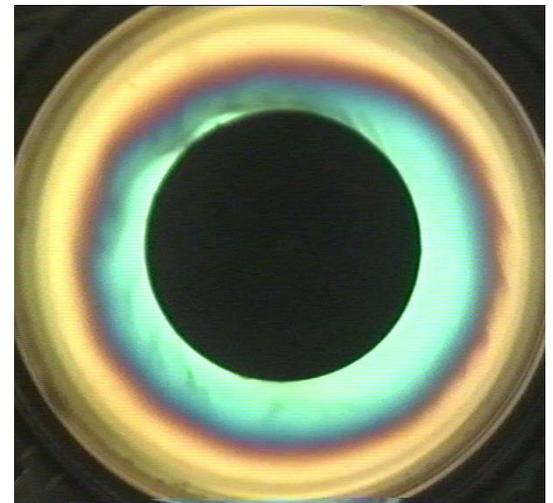
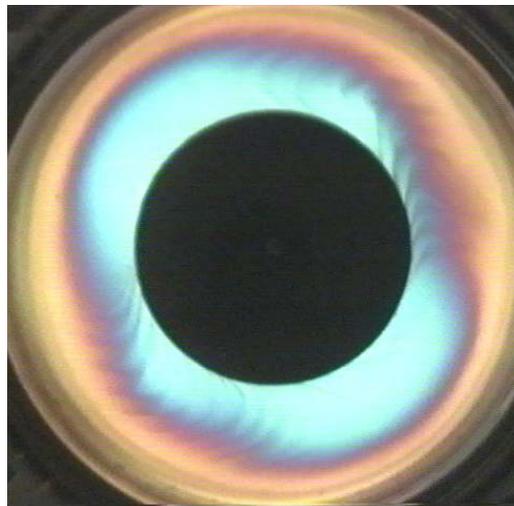
Impacts on the large-scale flow

# Gravity wave-induced flow transitions

without  
gravity  
waves



with gravity  
waves  
(‘noise’-  
induced  
transition)



# Crude stochastic gravity-wave parameterisation in QG equations:

$$\frac{\partial q_1}{\partial t} = \underbrace{\frac{1}{r} \frac{\partial \psi_1}{\partial \theta} \frac{\partial q_1}{\partial r} - \frac{1}{r} \frac{\partial \psi_1}{\partial r} \frac{\partial q_1}{\partial \theta}}_{\text{advection}} - \underbrace{\frac{\sqrt{\Omega \nu_1}}{H} [\nabla^2 \psi_1 + \chi_2 \nabla^2 (\psi_1 - \psi_2)]}_{\text{Ekman layer dissipation}} + \underbrace{\frac{2 \Delta \Omega \sqrt{\Omega \nu_1}}{H}}_{\text{lid}} + \eta$$

$$\frac{\partial q_2}{\partial t} = \underbrace{\frac{1}{r} \frac{\partial \psi_2}{\partial \theta} \frac{\partial q_2}{\partial r} - \frac{1}{r} \frac{\partial \psi_2}{\partial r} \frac{\partial q_2}{\partial \theta}}_{\text{advection}} - \underbrace{\frac{\sqrt{\Omega \nu_2}}{H} [\nabla^2 \psi_2 + \chi_1 \nabla^2 (\psi_2 - \psi_1)]}_{\text{Ekman layer dissipation}} - \eta$$

advection

Ekman layer dissipation

lid

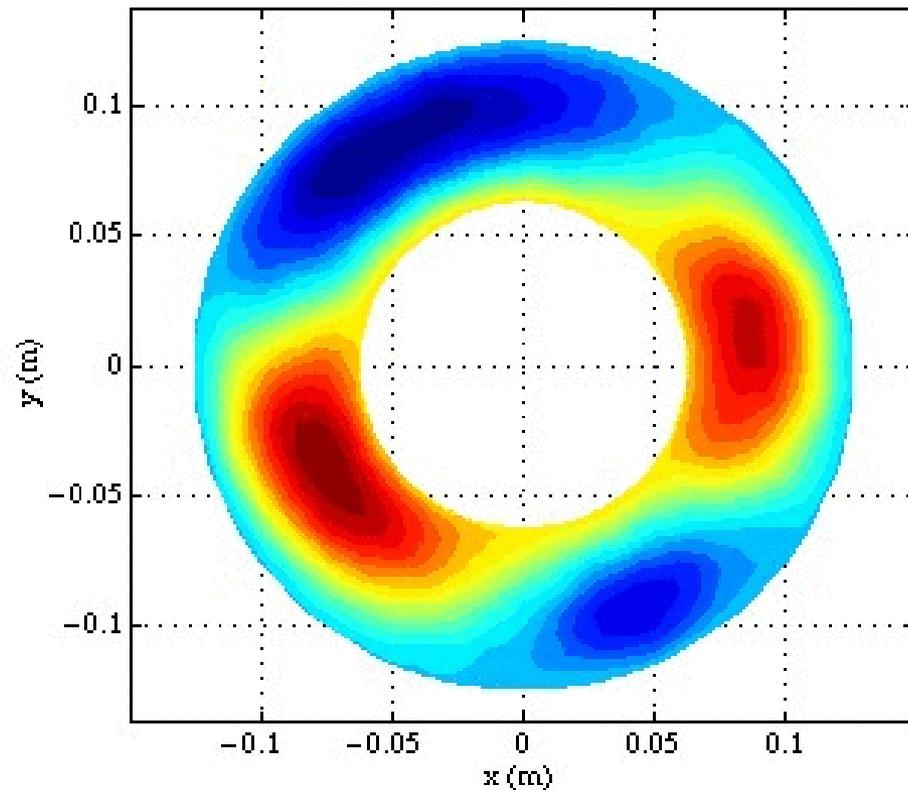
additive,  
uniformly-  
distributed  
white noise  
(purely  
baroclinic)

where:

$$q_1 = \nabla^2 \psi_1 + \frac{f^2}{g'H} (\psi_2 - \psi_1) + \frac{f}{H} \frac{r^2 \Omega^2}{2g}$$

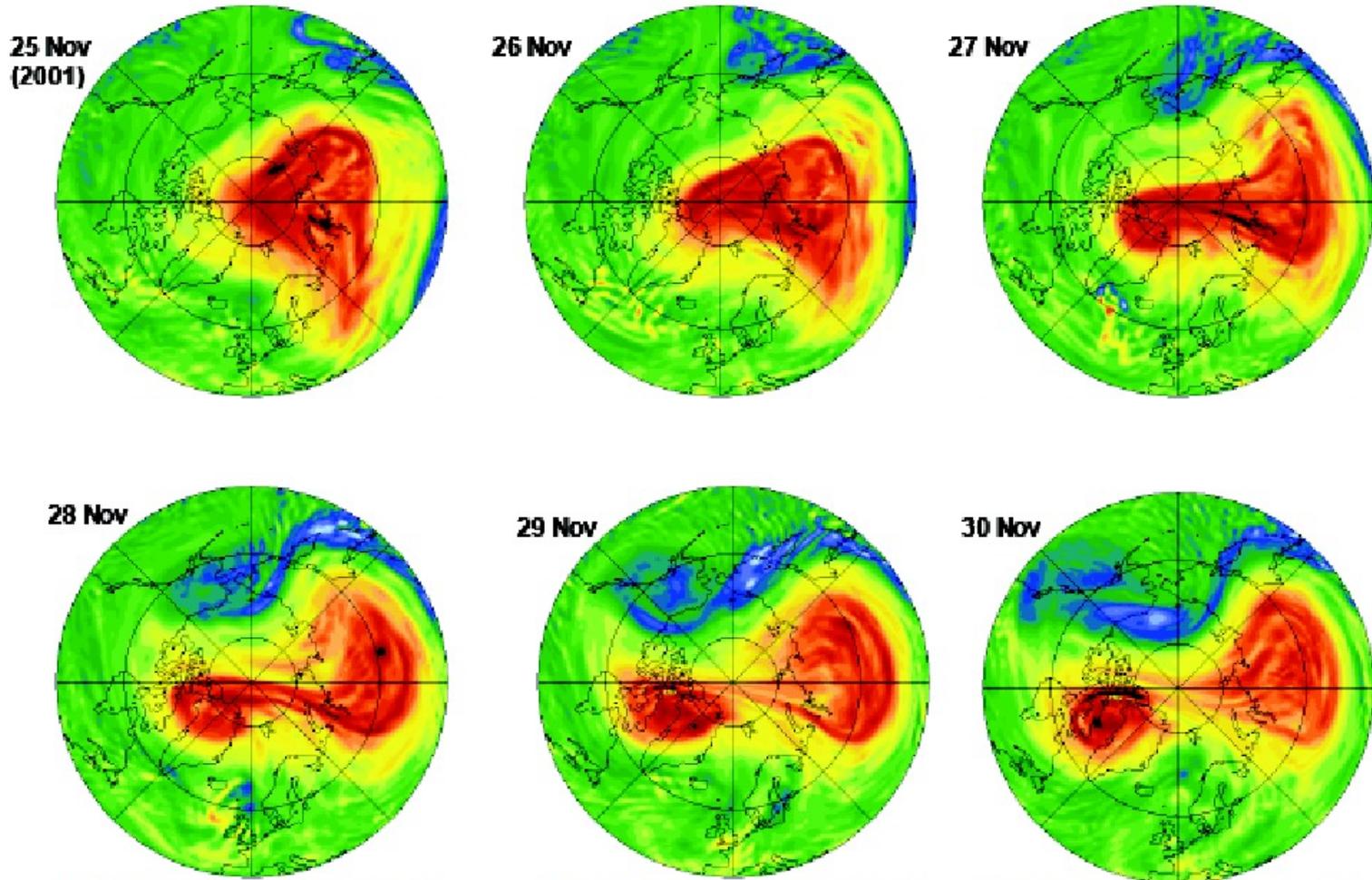
$$q_2 = \nabla^2 \psi_2 - \frac{f^2}{g'H} (\psi_2 - \psi_1) - \frac{f}{H} \frac{r^2 \Omega^2}{2g}$$

# Noise-induced transition in the *QUAGMIRE* quasi-geostrophic model



(Williams et al. 2004)

# Polar vortex splits as gravity wave noise-induced transitions



# Summary

- We appear to have observed the **spontaneous generation** of gravity waves by balanced flow in the laboratory
- The gravity wave amplitude varies **linearly** with  $Ro$ ... does anyone have an explanation for this?!
- **About 1%** of the balanced flow energy is lost to gravity waves each 'day'
- The gravity waves influence **large-scale flow transitions** and may be modelled as **random stochastic noise**

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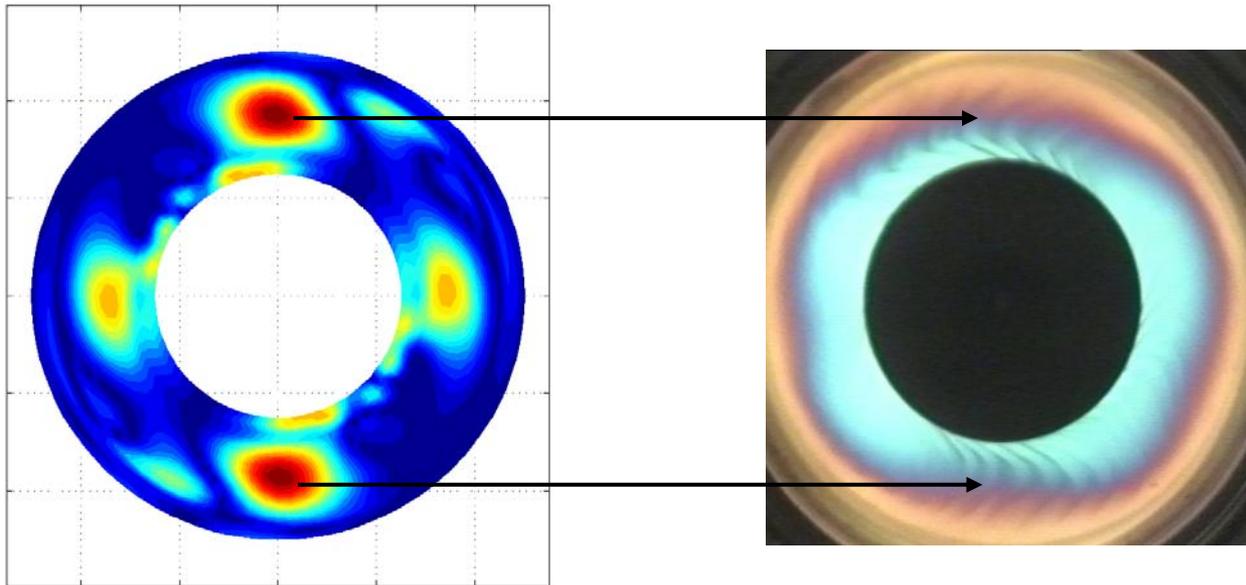
[www.met.reading.ac.uk/~williams](http://www.met.reading.ac.uk/~williams)

# Spontaneous-adjustment emission?

RSWEs (Ford, 1994): 
$$\left( \frac{\partial^2}{\partial t^2} + f^2 - gH\nabla^2 \right) \frac{\partial h}{\partial t} = \frac{\partial}{\partial t} \nabla \cdot \mathbf{G} + f \mathbf{k} \cdot \nabla \times \mathbf{G} + \frac{g}{2} \frac{\partial}{\partial t} \nabla^2 h^2$$

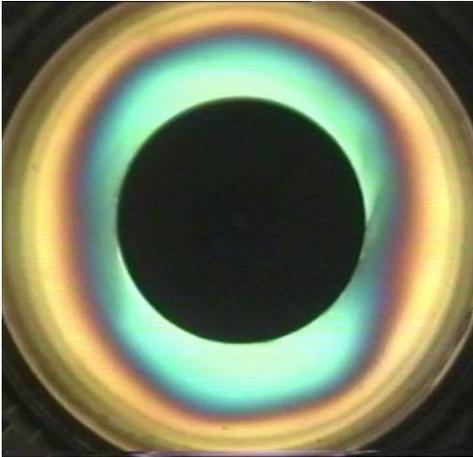
eigenmodes are GWs:  
$$\omega^2 = f^2 + gHK^2$$

GW source term, where  
$$\mathbf{G} = \mathbf{u} \nabla \cdot (h\mathbf{u}) + (h\mathbf{u} \cdot \nabla)\mathbf{u}$$

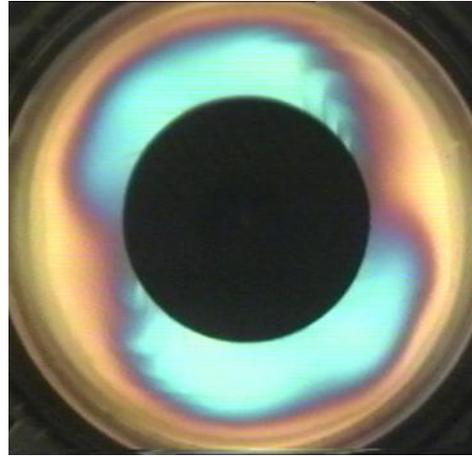


GW source term

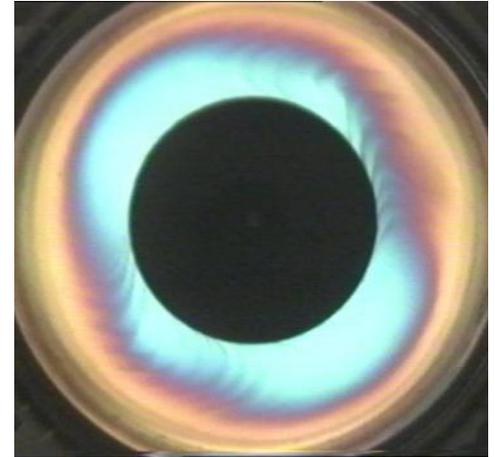
# Baroclinic lifecycle ( $\Delta t = 22$ s):



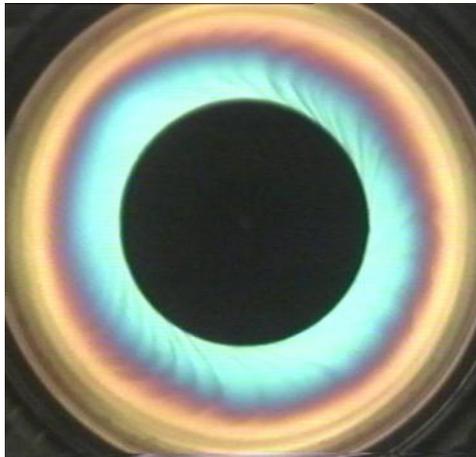
$\varphi = 0^\circ$



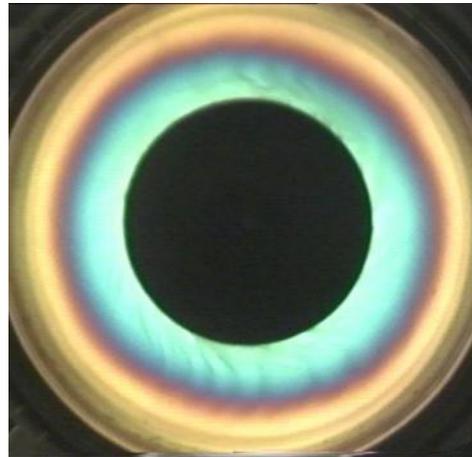
$\varphi = 60^\circ$



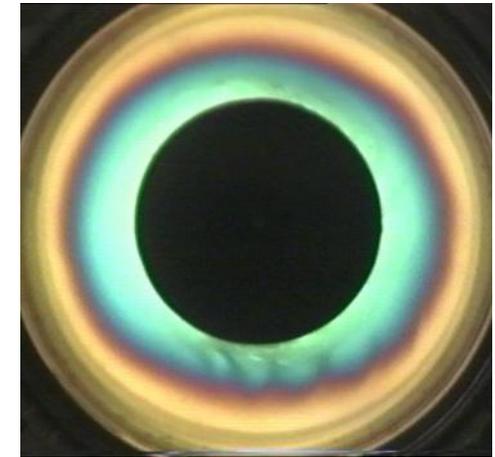
$\varphi = 120^\circ$



$\varphi = 180^\circ$

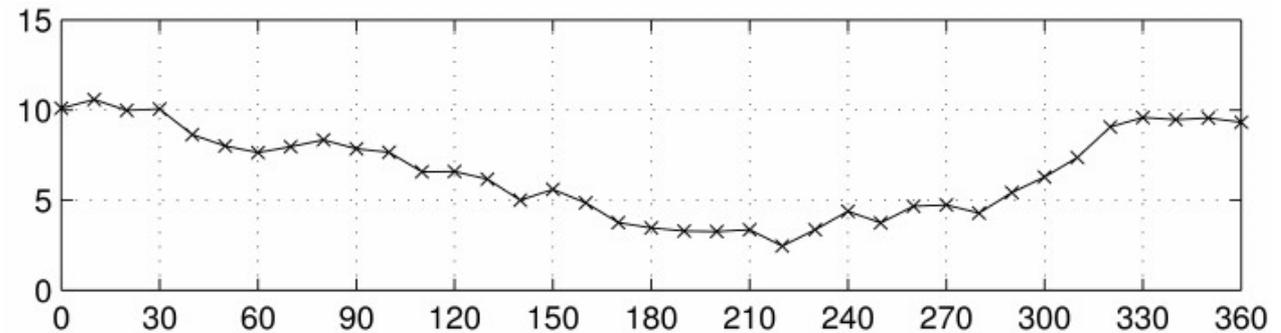


$\varphi = 240^\circ$

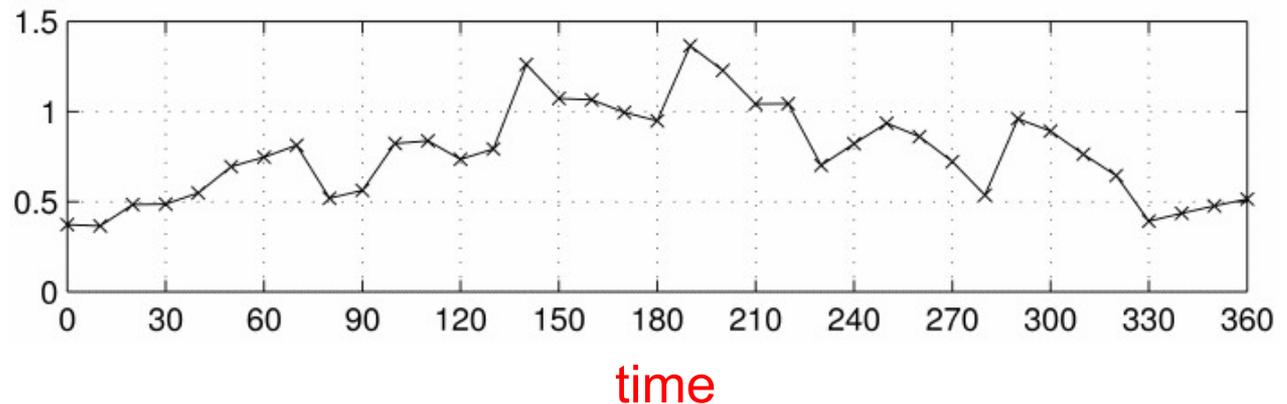


$\varphi = 300^\circ$   
(Williams et al. 2008)

large-scale flow  
amplitude (mm)



gravity wave  
amplitude (mm)



→ rate of energy growth in the gravity waves,  
as a fraction of the large-scale wave energy,  
is **~1% per turntable rotation period**

(Williams et al. 2008)